



TA 2-3

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## **Spectroscopic Observations of Novae V1065 CENTAURI and V1280 SCORPII using 45cm Cassegrain Telescope at Arthur C Clarke Institute, Sri Lanka**

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The spectroscopic observations of two novae namely nova Centauri 2007 (V1065 CEN) and nova Scorpii 2007 (V1280 SCO) were made by 45 cm Cassegrain telescope with  $f/12$  at Arthur C Clarke Institute, Sri Lanka during the period at 31st January to 20th Feb 2007. High resolution ( $\lambda/\Delta\lambda=22000$ ) profiles in  $H\alpha$  (6563 oA) region were obtain for V1065 CEN, 6, 15 and 20 days after maximum and  $H\alpha$  profiles of the same resolution were obtained for V1280 SCO, 4 days after maximum, covering the early decline stages of novae.

V1065 CEN is He/N-type spectra which characterize a broad (Gaussian FWHM 49 oA), saddle shaped and asymmetric  $H\alpha$  emission line with out prominent P-Cyg absorption component. Completely different  $H\alpha$  profile of V1280 SCO shows prominent P-Cyg absorption and narrow emission line (Gaussian FWHM 26 oA) which can be classified as Fe II type nova. The absence of prominent P-Cyg structure in V1065 CEN suggests that the emission causes by discrete shell while the prominent P-Cyg structure in V1280 SCO evidences a wind-like structure. The expansion velocities of these two systems measured from the minima of the P-Cyg profiles are close to 2300 km/s for V1065 CEN, 6 days after the maximum and 716 km/s for V1280 SCO, 4 days after the maximum.

The light curves V-t, B-t and visual-t have been used to estimate the distances of both novae. Based on the photometric analysis, the Nova V1065 CEN can be classified as fast ( $11 < t_2 < 25$ ) nova with the parameters  $t_{2V}=21$  days,  $t_{3V}=28$  days and  $t_{2B}=23$  days,  $t_{3B}=31$  days. The derived absolute magnitudes at maximum for nova V1065 CEN to be  $M_{0,V} = -7.58 \pm 0.18$  and  $M_{0,B} = -7.75 \pm 0.25$ . The mean distance module 16.57 and the color excess  $EB-V = +0.6$  correspond to a distance  $8.51 \pm 0.33$  kpc. The parameters  $t_{2V}=12$  days and  $t_{3V}=14$  days were calculated from visual-t light curve for nova V1280 SCO and It can be determine that the nova is in between very fast and fast nova. The rate of decline at  $t_2$ , 0.48 mag/d (very fast  $> 0.2$  mag/d) clearly determine that V1280 SCO is classified as very fast nova. The mean absolute magnitude at maximum is calculated to be  $M_{0,V} = -8.7 \pm 0.1$ . Neglecting the interstellar reddening the estimated distance to the nova V1280 SCO is  $3.2 \pm 0.2$  kpc.